
On the computation of apparent direct solar radiation

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Near-forward-scattered radiation coming from the vicinity of Sun's direction impacts the interpretation of measurements of direct solar radiation by pyrheliometers and sun photometers: a substantial part of the measured "direct" radiation can actually be scattered radiation. This issue is also relevant for concentrating solar power applications. Here, a Monte Carlo radiative transfer model is employed to study the apparent direct solar transmittance $t(\alpha)$, that is, the transmittance measured by an instrument that receives the radiation within a half-field-of-view (half-FOV) angle α from the centre of the solar disk, for various ice cloud, water cloud and aerosol cases. The contribution of scattered radiation to $t(\alpha)$ increases with increasing particle size, and it also depends strongly on ice crystal shape and roughness. The Monte Carlo calculations are compared with a simple approach, in which $t(\alpha)$ is estimated through Beer's law, using a scaled optical depth that excludes the part of the phase function corresponding to scattering angles smaller than α . Overall, this optical depth scaling approach works very well, although with some degradation of the performance for ice clouds for very small half-FOV angles ($\alpha < 0.5 - 1$ deg), and in optically thick cases. The errors can be reduced by fine-tuning the optical depth scaling factors based on the Monte Carlo results. It is also shown that the optical depth scaling used in delta-two-stream approximations employed in numerical weather prediction and climate model radiation calculations is inappropriate for simulating the direct solar radiation received by pyrheliometers.