#### Astrophysical light scattering problems (PAP316) Lecture 2c

## Microwave scattering in planetary science

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#### Radar echo from planetary surfaces





## "OC" and "SC" polarizations

#### Smooth surfaces:

Specular reflection -> All echo in the oppositecircular (OC) polarization than the transmitted signal

 Rough surfaces (wavelengthscale surface roughness or boulders): Quasi-specular + diffuse scattering

 > Echo partly in the OC polarization and partly in the same-circular (SC) polarization

#### **Radar cross section:**

$$\sigma_{Pol} = \frac{4\pi R^4 \lambda^2 P_{rx,Pol}}{P_{tx} A_{eff}^2}$$

**Radar albedo:** 

$$\hat{\sigma}_{Pol} = \frac{\sigma_{Pol}}{A_{proj}}$$

#### One-parameter definitions for surface roughness

- Traditionally used parameters for surface roughness:
  - Circular-polarization ratio  $(\mu_C)$ , typical values 0-1.0
  - Backscatter gain factor (g), typical values 1.2-1.5
  - Surface slope parameter (*C* or *s*), typical values 0.1-0.8

$$u_C = \frac{\sigma_{SC}}{\sigma_{OC}}$$

$$g = \hat{\sigma}_{OC} / R_F$$

$$[g = 1 \text{ for a smooth surface}]$$

$$g \ge 1$$

$$R_F = \left| \frac{\sqrt{\varepsilon} - 1}{\sqrt{\varepsilon} + 1} \right|^2$$

$$[\varepsilon : \text{electric permittivity}]$$

#### CPR DEPENDS ON THE TAXONOMIC TYPE?



## Understanding radar scattering

- Multi-parameter problem: The size, shape, and material of wavelength-scale particles on/in asteroid surfaces play a role in radar scattering as well as the structure of the sub-surface
- Lunar radar analysis has a long history that can help us to understand also radar scattering in asteroid surfaces
- Multi-wavelength comparative analyses
   can provide constraints
- Modeling work is crucial for understanding how to interpret the radar data





## Radar scattering laws

On the left: Backscatter coefficient as a function of the incidence angle

Several radar scattering laws have been developed (with different conditions)



Also: Considering only surface undulations is not enough!



## m-chi decomposition

Red = 
$$[mS_1(1 + \sin 2\chi)/2]^{1/2}$$
  $m = \frac{\sqrt{S_2^2 + S_3^2 + S_4^2}}{S_1}$  Degree of polarization  
Green =  $[S_1(1 - m)]^{1/2}$   
Blue =  $[mS_1(1 - \sin 2\chi)/2]^{1/2}$   $\chi = \frac{1}{2} \arcsin\left(\frac{S_4}{mS_1}\right)$  Degree of ellipticity  
 $M$  (Raney et al. 2012, JGR 17]

We denote the ratio of HH- to VVpolarized backscattered power for a dihedral pair, or suite of such features, as  $\alpha_{\rm D}$ . For an ideal dihedral feature,  $\sigma_{\rm HV}^0$  is negligible, and  $\beta$  is -1. Under these conditions,

$$Re[S_{HH}S_{VV}^*] = -\sqrt{\sigma_{HH}^0 \sigma_{VV}^0} = -\sigma_{VV}^0 \sqrt{\alpha_D}$$
(14)

and the CPR can be simplified from equations (7) and (8):

$$\mu_c = \frac{\left(1 + \sqrt{\alpha_D}\right)^2}{\left(1 - \sqrt{\alpha_D}\right)^2} \tag{15}$$



Figure 7. Circular polarization ratio (solid curve) and  $\alpha_D$ , the ratio of HH-polarized to VV-polarized backscatter (dotted line), averaged over a distribution of randomly oriented dihedral facet pairs, as a function of the real dielectric constant of the facets.

## Is it double scattering?



**Figure 6.** Diagram showing scattering geometry for a corner reflector. The vertical dashed line is the normal to the background plane surface.

#### Campbell et al. (2012)

### 1999 JM8 (P type)



0.0

0.2 0.4 0.6 0.8 Normalized (Fractional) RGB Value (0-1) 1.0

Low CPR all over, OC (blue) dominates [Hickson et al. 2020]

## 33342 1998 WT24 (E type)



m-chi



High CPR all over, SC (red) dominates: due to the wavelength-scale particles?

[Hickson et al. 2020]



# M-chi decomposition of individual particles

#### Single scattering

#### **Double scattering**



#### Backscattering as a function of size parameter

m = 1.43

m = 2.54



## Backscattering as a function of size parameter and refractive index



