

## Astrofysiikan peruskurssi I – Kaavakokoelma

- Säteilyn perusmääritelmiä:

$$I_\nu(\theta, \phi) = \frac{dE_\nu}{dt \cos \theta dAd\nu d\omega}$$

$$F_\nu = \int_{\Omega} I_\nu(\theta, \phi) \cos \theta d\omega$$

$$J_\nu = \frac{1}{4\pi} \int_{\Omega} I_\nu(\theta, \phi) d\omega$$

$$H_\nu = \frac{1}{4\pi} \int_{\Omega} I_\nu(\theta, \phi) \cos \theta d\omega = \frac{F_\nu}{4\pi}$$

$$K_\nu = \frac{1}{4\pi} \int_{\Omega} I_\nu(\theta, \phi) \cos^2 \theta d\omega = \frac{c}{4\pi} P_R$$

$$u = \frac{1}{c} \int_{\Omega} I_\nu(\theta, \phi) d\omega$$

- Mustan kappaleen säteily:

$$I_\nu = B_\nu(T) = \frac{2h\nu^3}{c^2} \frac{1}{e^{h\nu/kT}-1}$$

$$\lambda_{\max} T = 0.289782 \text{ cmK}$$

$$F_{\text{tot}} = \pi \int_0^\infty B_\nu d\nu = \sigma T_{\text{eff}}^4, \quad \sigma = 5.669 \cdot 10^{-8} \text{ Wm}^{-2}\text{K}^{-4}$$

- Säteilyn emissio ja absorptio:

$$j_\nu = \frac{1}{4\pi} \frac{dE_\nu}{dtd\nu dm}$$

$$k_\nu = \frac{\kappa_\nu}{\rho}$$

$$\tau_\nu(x) = \int_0^x \rho(x') k_\nu(x') dx'$$

$$\text{Kirchoffin laki: } \frac{j_\nu}{k_\nu} = B_\nu(T)$$

- Säteilykuljetus:

$$\cos \theta \frac{dI_\nu(\tau_\nu, \theta)}{d\tau_\nu} = I_\nu(\tau_\nu, \theta) - S_\nu(\tau_\nu), \quad S_\nu = \frac{j_\nu}{k_\nu}$$

$$\cos \theta \frac{dI_\nu}{d\tau_\nu} = I_\nu - \left( \frac{k_\nu}{k_\nu + \sigma_\nu} B_\nu + \frac{\sigma_\nu}{k_\nu + \sigma_\nu} J_\nu \right)$$

$$I_\nu(\tau_\nu, \theta) = \int_{\tau_\nu}^\infty S_\nu(\tau'_\nu) e^{-(\tau'_\nu - \tau_\nu) \sec \theta} \sec \theta d\tau'_\nu, \quad \sec \theta = 1/\cos \theta$$

$$J_\nu(\tau_\nu) = \frac{1}{2} \int_0^\infty S_\nu(\tau'_\nu) E_1(|\tau'_\nu - \tau_\nu|) d\tau'_\nu$$

$$F_\nu(\tau_\nu) = 2\pi \int_{\tau_\nu}^\infty S_\nu(\tau'_\nu) E_2(\tau'_\nu - \tau_\nu) d\tau'_\nu - 2\pi \int_0^{\tau_\nu} S_\nu(\tau'_\nu) E_2(\tau_\nu - \tau'_\nu) d\tau'_\nu$$

$$\text{Integraaliesponenttifunktio: } E_n(x) = \int_1^\infty \frac{e^{-xy}}{y^n} dy$$

- Kaasumaisen tilan fysiikkaa:

$$PV = NRT, \quad P = NkT, \quad P = \frac{\rho kT}{m}$$

$$P = \frac{1}{3}m\overline{v^2}, \quad \frac{1}{2}m\overline{v^2} = \frac{3}{2}kT$$

Adiabaattinen muutos:  $PV^\gamma = \text{vakio}$ ,  $TV^{\gamma-1} = \text{vakio}$ ,  $P^{1-\gamma}T^\gamma = \text{vakio}$

- Maxwellin nopeusjakautuma, Boltzmannin virityskaava sekä Ionisaatioyhtälö:

$$dN(v_x, v_y, v_z) = N \left( \frac{m}{2\pi kT} \right)^{3/2} e^{-\frac{m}{2kT}(v_x^2 + v_y^2 + v_z^2)} dv_x dv_y dv_z$$

$$dN(v) = 4\pi v^2 N \left( \frac{m}{2\pi kT} \right)^{3/2} e^{-\frac{m}{2kT}v^2} dv$$

$$\frac{N_n}{N_m} = \frac{g_n}{g_m} e^{-(E_n - E_m)/kT}$$

$$\frac{N_n}{N} = \frac{g_n}{u(T)} e^{-\chi_n/kT}$$

$$\frac{N_{i+1}N_e}{N_i} = \frac{(2\pi mkT)^{3/2}}{h^3} \frac{2u_{i+1}(T)}{u_i(T)} e^{-\chi_i/kT}$$

$$\frac{N_{i+1}P_e}{N_i} = 0.331 T^{5/2} \cdot \frac{2u_{i+1}(T)}{u_i(T)} 10^{-\frac{5040}{T}\chi_i},$$

$[P_e] = \text{dyn/cm}^2$ ,  $[T] = \text{K}$ ,  $[\chi_i] = \text{eV}$