

Taylor's conjecture on magnetic helicity conservation in magnetohydrodynamics

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Incompressible **magnetohydrodynamics (MHD)** is a continuum model for electrically conducting fluids.

In order to formulate the equations, we start with

$$\rho(\partial_t + \mathbf{u} \cdot \nabla)\mathbf{u} = -\nabla p + \nu \Delta \mathbf{u} + \mathbf{J} \times \mathbf{B}, \quad (1)$$

$$\nabla \cdot \mathbf{u} = 0, \quad (2)$$

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{J}, \quad (3)$$

$$\partial_t \mathbf{B} + \nabla \times \mathbf{E} = \mathbf{0}, \quad (4)$$

$$\nabla \cdot \mathbf{B} = 0, \quad (5)$$

where $\nu > 0$. (We set $\rho = \mu_0 = 1$ for notational simplicity.)

By combining (3)–(4) with Ohm's law $\mathbf{E} = \mathbf{B} \times \mathbf{u} + \eta \mathbf{J}$ we get the induction equation

$$\partial_t \mathbf{B} + \nabla \times (\mathbf{B} \times \mathbf{u}) + \eta \nabla \times (\nabla \times \mathbf{B}) = \mathbf{0}.$$

We eliminate \mathbf{E} and \mathbf{J} altogether and formulate the MHD equations in \mathbf{B} , \mathbf{u} and p :

The **(homogeneous,) incompressible, viscous, resistive MHD equations** consist of

$$\partial_t \mathbf{u} + (\mathbf{u} \cdot \nabla) \mathbf{u} = -\nabla p + \nu \Delta \mathbf{u} + (\nabla \times \mathbf{B}) \times \mathbf{B}, \quad (6)$$

$$\nabla \cdot \mathbf{u} = 0, \quad (7)$$

$$\partial_t \mathbf{B} + \nabla \times (\mathbf{B} \times \mathbf{u}) + \eta \nabla \times (\nabla \times \mathbf{B}) = 0, \quad (8)$$

$$\nabla \cdot \mathbf{B} = 0, \quad (9)$$

where $\nu, \eta > 0$.

When $\nu = \eta = 0$, (6)–(9) is called **ideal MHD**.

Given a bounded, simply connected domain $\mathcal{V} \subset \mathbb{R}^3$ with $\partial \mathcal{V} = \mathcal{S}$ we set the boundary conditions $\mathbf{u}|_{\mathcal{S}} = 0$; $\mathbf{B} \cdot \mathbf{n}|_{\mathcal{S}} = 0$; $\mathbf{E} \times \mathbf{n}|_{\mathcal{S}} = 0$, which leads to $(\nabla \times \mathbf{B}) \times \mathbf{n}|_{\mathcal{S}} = 0$.

In ideal MHD, we set $\mathbf{u} \cdot \mathbf{n}|_{\mathcal{S}} = \mathbf{B} \cdot \mathbf{n}|_{\mathcal{S}} = 0$.

Conserved integral quantities of ideal MHD

Continuously differentiable solutions \mathbf{u} , \mathbf{B} of *ideal* MHD conserve

$$\int_{\mathcal{V}} \frac{u^2 + B^2}{2} dV := \int_{\mathcal{V}} \frac{|\mathbf{u}(\mathbf{x}, t)|^2 + |\mathbf{B}(\mathbf{x}, t)|^2}{2} dV \quad (\text{total energy}),$$

$$\int_{\mathcal{V}} \mathbf{u} \cdot \mathbf{B} dV \quad (\text{cross helicity}),$$

$$\int_{\mathcal{V}} \mathbf{A} \cdot \mathbf{B} dV \quad (\text{magnetic helicity}),$$

where $\nabla \times \mathbf{A} = \mathbf{B}$.

However, simulations point towards **anomalous dissipation** of total energy: when viscosity and resistivity tend to zero, the energy dissipation rate tends to a *positive* constant, see

- Mininni-Pouquet (Phys. Rev. Lett. 2009),
- Dallas-Alexakis (Astrophys. J. Lett. 2014),
- Linkmann-Berera-McComb-McKay (Phys. Rev. E 2015).

Taylor (Phys. Rev. Lett. 1974) conjectured that magnetic helicity is, by contrast, approximately conserved at very low resistivities.

- Magnetic helicity was first studied by Woltjer (PNAS 1958).
- It is *gauge invariant* (independent of the choice of \mathbf{A}) in simply connected domains \mathcal{V} (when $\mathbf{B} \cdot \mathbf{n}|_{\partial\mathcal{V}} = 0$):

If $\nabla \times \mathbf{A} = \nabla \times \mathbf{A}' = \mathbf{B}$, then $\mathbf{A} - \mathbf{A}' = \nabla\phi$ and

$$\int_{\mathcal{V}} (\mathbf{A} - \mathbf{A}') \cdot \mathbf{B} \, dV = \int_{\mathcal{V}} \nabla\phi \cdot \mathbf{B} \, dV = \int_{\mathcal{V}} \nabla \cdot (\phi\mathbf{B}) \, dV = \int_{\mathcal{S}} \underbrace{\phi \mathbf{B} \cdot \mathbf{n}}_{=0} \, dS = 0.$$

- It measures e.g. the linkage of magnetic field lines, see Moffatt (J. Fluid Mech. 1969).
- For a review see Berger (Plasma Phys. Control. Fusion 1999).

Force-free magnetic fields

- In both astrophysical and laboratory experiments, plasmas tend to evolve towards a (Beltrami) state

$$\nabla \times \mathbf{B}(\mathbf{x}) = \lambda \mathbf{B}(\mathbf{x}) \quad (\lambda \text{ constant}), \quad (10)$$

see e.g. Qin-Liu-Li-Squire (Phys. Rev. Lett. 2012).

- First, Lüst and Schlüter (Z. Astrophys. 1954) and Woltjer (B.A.N. 1958) suggested that various cosmic magnetic fields are **force-free**, i.e., the Lorentz force $(\nabla \times \mathbf{B}) \times \mathbf{B} = 0$; then $\nabla \times \mathbf{B} = \lambda \mathbf{B}$, but λ could be non-constant.
- Woltjer (PNAS 1958) showed that magnetic helicity $\int_V \mathbf{A} \cdot \mathbf{B} dV$ is conserved in time by (smooth) solutions of ideal MHD.
- Woltjer noted that by minimizing the magnetic energy $\int_V B^2/2$ under the constraint $\int_V \mathbf{A} \cdot \mathbf{B} dV = K$ one obtains (10).
- Woltjer suggested that the plasma relaxes towards a state that minimises magnetic energy subject to the (fixed) magnetic helicity.

(Too) many conserved quantities of ideal MHD

- In ideal MHD, magnetic helicity is conserved for all volumes $\mathcal{V}_i(t) \subset \mathcal{V}$ moving with the fluid that are magnetically closed at $t = 0$ (i.e. $\mathbf{B}(\mathbf{x}, 0) \cdot \mathbf{n}(\mathbf{x})|_{\partial\mathcal{V}_i(0)} = 0$).
- This can lead to numerous conserved quantities $K_i := \int_{\mathcal{V}_i(t)} \mathbf{A}(\mathbf{x}, t) \cdot \mathbf{B}(\mathbf{x}, t) dV$, sometimes called *subhelicities*.
- Formally, if one minimises magnetic energy subject to all the constraints $\int_{\mathcal{V}_i} \mathbf{A} \cdot \mathbf{B} dV = K_i$, one gets $\nabla \times \mathbf{B}(\mathbf{x}, t) = \lambda_i(t)\mathbf{B}(\mathbf{x}, t)$ in each $\mathcal{V}_i(t)$. Here $\lambda_i(t)$ is directly related to the initial conditions.
- However, a relaxed state is observed to be essentially independent of initial conditions! (See Ortolani-Schnack: Magnetohydrodynamics of Plasma Relaxation, 1993)

- Taylor (Phys. Rev. Lett. 1974) argued that with non-zero resistivity η , the subhelicities $K_i = \int_{\mathcal{V}} \mathbf{A} \cdot \mathbf{B} dV$ would no longer be conserved but $K_0 = \int_{\mathcal{V}} \mathbf{A} \cdot \mathbf{B} dV$ itself would remain a good invariant.
- Thus, when $\eta > 0$, one should minimise the magnetic energy subject to the constraint that $K_0 = \int_{\mathcal{V}} \mathbf{A} \cdot \mathbf{B} dV$ remain constant.
- It is then hypothesised that during an initial turbulent state, magnetic fields in various laboratory settings shed extra magnetic energy until they reach a quiescent 'relaxed state' where $\nabla \times \mathbf{B} = \lambda \mathbf{B}$.
- The resulting **Taylor relaxation theory** is discussed in Ortolani-Schnack: Magnetohydrodynamics of Plasma Relaxation, 1993.
- However, Taylor's theory does not address the dynamics that lead to the relaxed state $\nabla \times \mathbf{B} = \lambda \mathbf{B}$. A dynamical explanation was proposed in Qin-Liu-Li-Squire (Phys. Rev. Lett. 2012).

- **Taylor's conjecture (1974):** (when \mathcal{V} is simply connected and $\mathbf{B} \cdot \mathbf{n}|_S = 0$), magnetic helicity is approximately conserved in \mathcal{V} for very small resistivities $\eta > 0$.
- When $\nu, \eta > 0$,

$$\begin{aligned}\partial_t \int_{\mathcal{V}} \mathbf{A}(x, t) \cdot \mathbf{B}(x, t) dV &= -2\eta \int_{\mathcal{V}} \mathbf{B}(x, t) \cdot \nabla \times \mathbf{B}(x, t) dV, \\ \partial_t \int_{\mathcal{V}} (B^2 + u^2)/2 dV &= -\nu \int_{\mathcal{V}} |\nabla \times \mathbf{u}|^2 dV - \eta \int_{\mathcal{V}} |\nabla \times \mathbf{B}|^2 dV.\end{aligned}$$

- For small resistivities $\eta > 0$, Berger (Geophys. Astrophys. Fluid Dynamics 1984) showed (under certain extra physical hypotheses) that magnetic helicity dissipates much slower than magnetic energy.
- Mathematical version, see Caglisch-Klapper-Steele (Comm. Math. Phys. 1997):

Conjecture (Taylor 1974)

Magnetic helicity does not dissipate in the inviscid, non-resistive limit.

- Given smooth initial data \mathbf{u}_0 and \mathbf{B}_0 , it is wide open whether the Cauchy problem for MHD has a smooth solution.
- However, a weak 'Leray-Hopf solution' exists, see Sermange-Temam (Commun. Pure Appl. Math. 1984).
- Denote

$$L^2_\sigma(\mathcal{V}, \mathbb{R}^3) := \{\mathbf{v} : \int_{\mathcal{V}} v^2 dV < \infty, \nabla \cdot \mathbf{v} = 0, \mathbf{v} \cdot \mathbf{n}|_{\partial\mathcal{V}} = 0\}.$$

Definition

Let $\mathbf{u}_0, \mathbf{B}_0 \in L^2_\sigma(\mathcal{V}, \mathbb{R}^3)$. Then (\mathbf{u}, \mathbf{B}) is called a **Leray-Hopf solution** if

- \mathbf{u} and \mathbf{B} solve viscous, resistive MHD with initial data $\mathbf{u}_0, \mathbf{B}_0$,
- $\mathbf{u}, \mathbf{B} \in L^\infty(0, T; L^2) \cap L^2(0, T; W^{1,2})$,
- \mathbf{u} and \mathbf{B} satisfy, at every $t \in (0, T)$, the *energy inequality*

$$\begin{aligned} & \frac{1}{2} \int_{\mathcal{V}} (u^2 + B^2) dV \\ \leq & \frac{1}{2} \int_{\mathcal{V}} (u_0^2 + b_0^2) dV - \int_0^t \int_{\mathcal{V}} (\nu |\nabla \times \mathbf{u}|^2 + \eta |\nabla \times \mathbf{B}|^2) dV d\tau. \end{aligned}$$

Theorem (Faraco-L. (to appear in Comm. Math. Phys.))

Suppose

- $\nu_j, \eta_j \searrow 0$ when $j \rightarrow \infty$.
- At each j , $(\mathbf{u}_j, \mathbf{B}_j)$ is a Leray-Hopf solution with initial data $(\mathbf{u}_{0,j}, \mathbf{B}_{0,j})$,
- $\mathbf{u}_j \rightharpoonup \mathbf{u}$, $\mathbf{B}_j \rightharpoonup \mathbf{B}$, $\mathbf{u}_{0,j} \rightarrow \mathbf{u}_0$, $\mathbf{B}_{0,j} \rightarrow \mathbf{B}_0$.

Then \mathbf{B} conserves magnetic helicity (a.e. in time).

Similar ideas had been used elsewhere:

- Cheskidov-Lopes Filho-Nussenzveig Lopes-Shvydkoy (Comm. Math. Phys. 2016): on \mathbb{T}^2 , if a solution \mathbf{u} of Euler equations is a weak inviscid limit of Leray-Hopf solutions, then (under reasonably mild assumptions on initial data,) \mathbf{u} conserves kinetic energy.
- Constantin-Ignatova-Nguyen 2018: in SQG (in bounded, smooth domains), weak inviscid limits of Leray-Hopf solutions conserve the Hamiltonian.

Mathematical results on magnetic helicity conservation in ideal MHD:

- $\mathbf{u} \in C([0, T]; B_{3,\infty}^{\alpha_1}(\mathbb{T}^3; \mathbb{R}^3))$ and $\mathbf{B} \in C([0, T]; B_{3,\infty}^{\alpha_2}(\mathbb{T}^3; \mathbb{R}^3))$, $\alpha_1 + 2\alpha_2 > 0$:
Caflisch-Klapper-Steele (Comm. Math. Phys. 1997)
- $\mathbf{u}, \mathbf{B} \in L^3([0, T]; L^3(\mathbb{T}^3, \mathbb{R}^3))$: Kang-Lee (Nonlinearity 2007) and Aluie
(Ph.D. dissertation 2009) with Eyink
- Not in general conserved by $\mathbf{u}, \mathbf{B} \in L^\infty(0, T; L^2(\mathbb{T}^3, \mathbb{R}^3))$:
Beekie-Buckmaster-Vicol (arXiv 2019)
- Weak L^2 -limits of solutions (and subsolutions) $\mathbf{u}, \mathbf{B} \in L^3([0, T]; L^3(\mathbb{T}^3, \mathbb{R}^3))$:
Faraco-L.-Székelyhidi (arXiv 2019)

Thank you for your attention!